

## Double layers in the downward current region of the aurora

R. E. Ergun<sup>1, 2</sup>, L. Andersson<sup>2</sup>, C. W. Carlson<sup>3</sup>, D. L. Newman<sup>4</sup>, and M. V. Goldman<sup>4</sup>

<sup>1</sup>Department of Astrophysical and Planetary Sciences, University of Colorado, Boulder, CO, USA

<sup>2</sup>Laboratory for Atmospheric and Space Physics, University of Colorado, Boulder, CO, USA

<sup>3</sup>Space Sciences Laboratory, University of California, Berkeley, CA, USA

<sup>4</sup>Center for Integrated Plasma Studies, University of Colorado, Boulder, CO, USA

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**Abstract.** Direct observations of magnetic-field-aligned (parallel) electric fields in the downward current region of the aurora provide decisive evidence of naturally occurring double layers. We report measurements of parallel electric fields, electron fluxes and ion fluxes related to double layers that are responsible for particle acceleration. The observations suggest that parallel electric fields organize into a structure of three distinct, narrowly-confined regions along the magnetic field ( $\mathbf{B}$ ). In the “ramp” region, the measured parallel electric field forms a nearly-monotonic potential ramp that is localized to  $\sim 10$  Debye lengths along  $\mathbf{B}$ . The ramp is moving parallel to  $\mathbf{B}$  at the ion acoustic speed ( $v_s$ ) and in the same direction as the accelerated electrons. On the high-potential side of the ramp, in the “beam” region, an unstable electron beam is seen for roughly another 10 Debye lengths along  $\mathbf{B}$ . The electron beam is rapidly stabilized by intense electrostatic waves and nonlinear structures interpreted as electron phase-space holes. The “wave” region is physically separated from the ramp by the beam region. Numerical simulations reproduce a similar ramp structure, beam region, electrostatic turbulence region and plasma characteristics as seen in the observations. These results suggest that large double layers can account for the parallel electric field in the downward current region and that intense electrostatic turbulence rapidly stabilizes the accelerated electron distributions. These results also demonstrate that parallel electric fields are directly associated with the generation of large-amplitude electron phase-space holes and plasma waves.

Measurements, strongly suggest that the electric field is self-consistently supported as a “strong” ( $e\Delta\Phi > T_{e,i}$ ) double layer (Block, 1972; Schamel and Bujarbarua, 1983). The above report (Ergun et al., 2001) also establishes a direct link between parallel electric fields and electron phase-space holes.

Most of the earlier work on parallel electric fields of the aurora focused on the upward current region. Visible auroral arcs were determined to result from energetic ( $\sim 10$  keV) electrons accelerated earthward by parallel electric fields (Evans, 1974; Shelley et al., 1976; Mozer et al., 1977). Theoretical treatments of the self-consistent structure include weak double layers (Temerin et al., 1982), strong double layers (Block, 1972), anomalous resistivity (Hudson and Mozer, 1978), and parallel electric fields associated with ion cyclotron waves (Ishiguro et al., 1997). However, no generally-accepted theoretical description of these parallel electric fields has emerged, primarily due to lack of detailed observations.

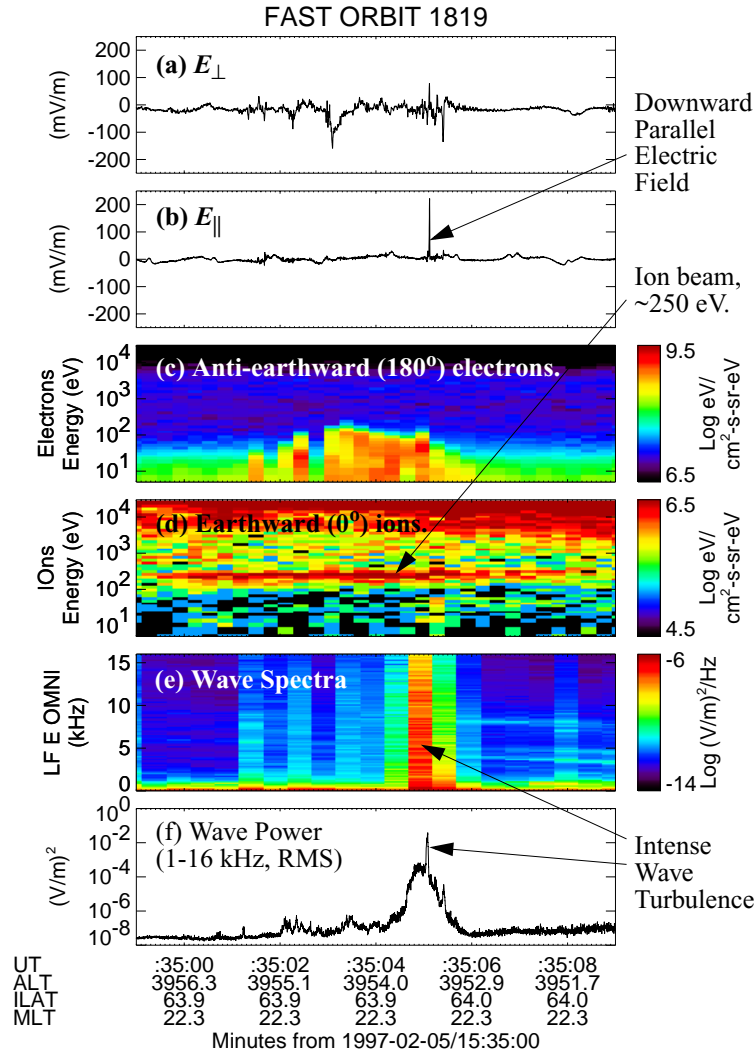
Direct observations of parallel electric fields in the upward current region have been reported (Mozer and Kletzing, 1998) but corroborative electron and ion fluxes are not available. The reported electric field amplitudes are greater than 100 mV/m and therefore suggest that the parallel electric field is localized. Collisionless, localized, parallel electric fields have also been extensively studied in laboratory experiments (Quon and Wong, 1976). Direct observations in the downward current region (Ergun et al., 2001) now indicate that the parallel electric fields are localized to roughly 10 Debye lengths ( $\lambda_D$ ) along  $\mathbf{B}$ .

Observations also suggest that the parallel electric field sets up strongly unstable electron beams which, in turn, create a spatially-separated region of strong electrostatic wave turbulence from which electron phase-space holes emerge (Matsumoto et al., 1994; Ergun et al., 1998a, b). The accelerated electron fluxes are substantially modified by the accompanying intense wave turbulence and electron phase-space holes. The wave turbulence modifies the electron distribution to the degree that part of the electron distribution is

### 1 Introduction

The purpose of this article is to present a detailed description of a direct observation of the parallel electric field in the downward current region of the aurora reported by Ergun et al. (2001). The earlier report included observations of electron fluxes which, together with the electric field mea-

Correspondence to: R. E. Ergun (ree@fast.colorado.edu)



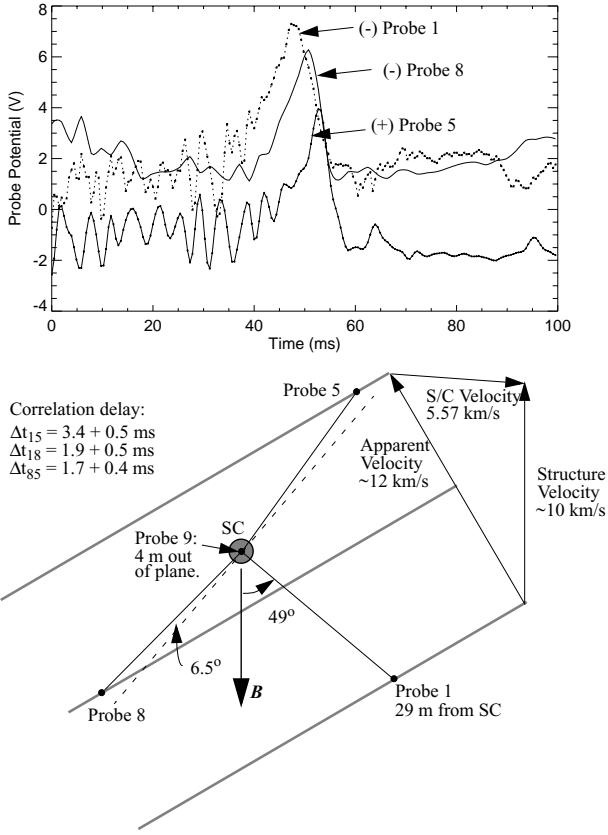
**Fig. 1.** Derived from Ergun et al. (2001) (a) The electric field perpendicular to  $\mathbf{B}$  ( $E_{\perp}$ ) and nearly along the velocity of the satellite. (b) The electric field parallel to  $\mathbf{B}$ . (c) Anti-earthward travelling ion energy flux versus energy and time. The ion fluxes have pitch angles from  $-22.5^{\circ}$  to  $22.5^{\circ}$ . (d) Earthward travelling electron energy flux versus energy and time. The electron fluxes have pitch angles from  $167.5^{\circ}$  to  $202.5^{\circ}$ . (e) The spectral power density of the omni-directional electric field from  $\sim 40$  Hz to 16 kHz. Each spectra is averaged over  $\sim 0.25$  s and has 40 Hz resolution in frequency. (f) The wave power integrated from  $\sim 1$  kHz to 16 kHz displayed at 4 ms resolution.

reflected back toward the double layer. These reflected electrons may provide a mechanism for interaction between the wave turbulence and the double layer. These data suggest that a self-consistent treatment of the parallel electric field must include substantial modification of the electron distribution due to intense wave turbulence.

## 2 Observations

The observations are from the Fast Auroral SnapshoT (FAST) satellite (Carlson et al., 1998a) which measures electromagnetic fields and charged particle distributions in the earth's auroral zone. The FAST satellite is in a near-polar orbit with  $83^{\circ}$  inclination. The orbit has a 350 km perigee and a 4175 km apogee. The FAST satellite has electron and ion spectrographs and 3-axis electric field and magnetic field instruments.

Figure 1 displays 10 s of data from the downward current region on the aurora. The spacecraft was in the pre-midnight sector travelling toward the northern polar cap at 5.57 km/s so the horizontal axis in Fig. 1 also may be interpreted as distance spanning approximately 56 km. The top two panels display the (a) perpendicular (to  $\mathbf{B}$ ) electric field ( $E_{\perp}$ ) closest to the direction of the spacecraft velocity and (b) the parallel electric field ( $E_{\parallel} = \mathbf{E} \cdot \mathbf{B}/|\mathbf{B}|$ ). Immediately below (Panel c) are the anti-earthward electron energy fluxes with pitch angles between  $157.5^{\circ}$  and  $202.5^{\circ}$ . The electron fluxes are plotted as an energy-time-flux spectrogram. The energy ranges from 5 eV to 30 keV. Panel (d) is an energy-time-flux spectrogram of the earthward-moving ion fluxes ( $-22.5^{\circ}$  to  $22.5^{\circ}$  pitch angles). Panels (e) and (f) display the omni-directional wave electric field in the frequency range from  $\sim 1$  kHz to 16 kHz. Panel (e) displays the spectral power density as a function of frequency (40 Hz resolution) and time (0.25 s res-



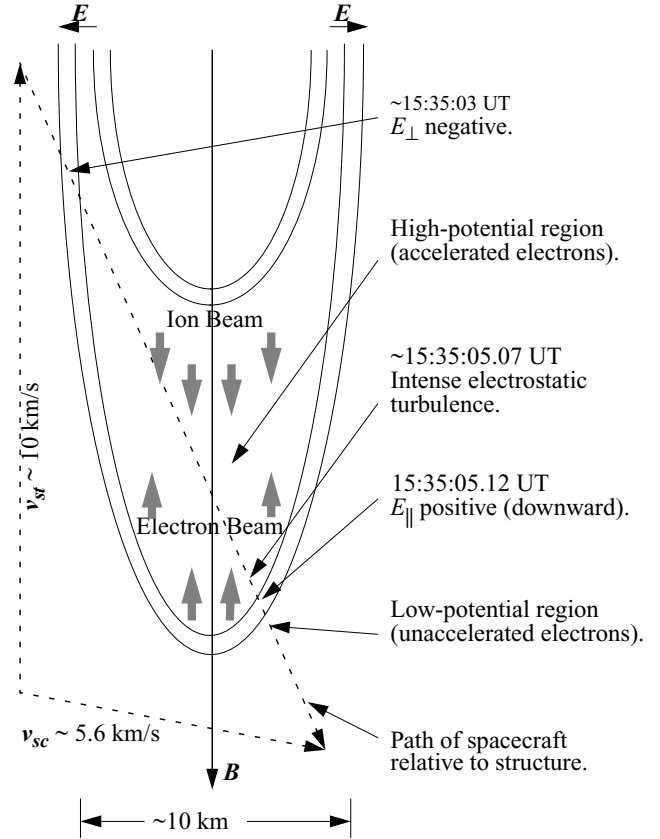
**Fig. 2.** (Top) The signals on three spatially separated probes during the  $E_{\parallel}$  event. The dotted line is the negative of the Probe 1 signal, the solid line is the negative of the Probe 8 signal, and the dotted solid line is the Probe 5 signal. (Bottom) The orientation of the spacecraft and the position of the probes during the  $E_{\parallel}$  event. The shaded lines are perpendicular to the motion of the structure. Probes 5 and 8 are  $6.5^{\circ}$  from perpendicular to Probe 1 as indicated by the dashed line.

olution). Panel (f) plots the RMS integrated wave power at 4 ms resolution.

At  $\sim 15:35:03$  UT,  $E_{\perp}$  reaches  $-150$  mV/m (Fig. 1a) and the 10 eV to 100 eV anti-earthward electron fluxes increase (Fig. 1c). These observations are characteristic of the downward current region of the aurora and indicate that a parallel electric field is on the magnetic flux tube and earthward of the spacecraft (Carlson et al., 1998b; Ergun et al., 1998c). The 200 mV/m (positive is earthward in the northern hemisphere) excursion in  $E_{\parallel}$  at  $\sim 15:35:05$  UT (Fig. 1b), however, is extremely unusual. At the same time, the electron fluxes abruptly decrease (Fig. 1c) and the wave power briefly increases by nearly two orders of magnitude as indicated by the spike in Fig. 1f and the broad band increase in wave spectral power density in Fig. 1e.

### 2.1 Electric field observations

One of the first problems that needs to be resolved is the spatial-temporal nature of the large-amplitude spike in  $E_{\parallel}$ .

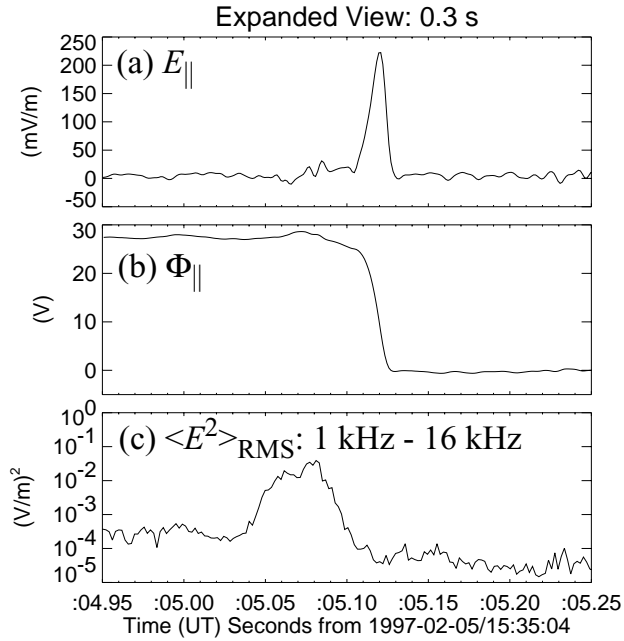


**Fig. 3.** A cartoon interpreting the observations in Fig. 1.

Figure 2 shows the signals from three individual probes and the magnetic orientation of spacecraft during the event. The magnetic field, pointing downward in the diagram, was within  $3^{\circ}$  of lying in the spin plane of the satellite. The spacecraft is moving at 5.57 km/s toward the right in the diagram; its velocity is almost perpendicular to  $B$  and downward.

The horizontal axis of the top panel is greatly magnified in time encompassing only 100 ms. The spike in  $E_{\parallel}$  first appears on Probe 1, next on Probe 8, the last on Probe 5. The delay between the peaks of Probe 1 and Probe 8 appear to be  $\sim 5$  ms while the time delay derived from correlating the two signals over a 20 ms period (40 ms to 60 ms in Fig. 2) is  $\sim 3.4$  ms. The direction of motion of the structure in the spacecraft frame is clearly from the bottom right toward the top left in the diagram. The velocity of the structure derived from the correlation delay is 12.5 m/s. In an inertial frame, the structure velocity is nearly anti-parallel to  $B$  at  $\sim 10$  km/s. Accounting for uncertainties in the measurements, the electric field observations are consistent with an  $E_{\parallel}$  structure travelling anti-earthward within  $20^{\circ}$  of  $B$  and with a velocity between 8 km/s and 11 km/s. This speed is consistent with the estimated ion acoustic speed on the low-potential side (see later). Using these speeds, the net potential of the  $E_{\parallel}$  structure is determined to be between 21 V and 30 V.

Since the  $E_{\parallel}$  measurement is critical to this article, we



**Fig. 4.** An expanded view of the  $E_{\parallel}$  event. (a)  $E_{\parallel}$ . (b)  $\int E_{\parallel} v_{\parallel} dt$ .  $v_{\parallel}$  is assumed to be 10 km/s. (c) The wave power integrated from  $\sim 1$  kHz to 16 kHz displayed at 4 ms resolution.

describe our methods of validating the observations. The FAST electric field instrument uses a set of four, 8-cm diameter, high impedance spherical probes to measure the vector field (Fig. 2). In the spin plane of the satellite, the electric field is derived from long-baseline (56 m and 29 m) dipoles which measure the instantaneous (independent of satellite spin phase) electric field. From empirical analysis, we find that the accuracy of the measurement depends on the plasma properties, in particular, the plasma density. With plasma density ( $n_e$ )  $> 5 \text{ cm}^{-3}$ , the spin-averaged accuracy is  $\pm 1 \text{ mV/m}$  and the instantaneous accuracy is  $\pm 10 \text{ mV/m}$  if (1) none of the probes are magnetically conjunctive to the spacecraft, (2) none of the probes is in dc saturation ( $\pm 50 \text{ V}$ ), (3) none of the probes is in ac saturation ( $\pm 300 \text{ mV/m}$ ), (4) no bias or mode switching was occurring at the time, and (5) the spacecraft was not crossing the terminator. Such conditions were met during the event.

The spin-axis component of  $\mathbf{E}$  is measured with far less accuracy ( $\pm 200 \text{ mV/m}$  with  $n_e > 5 \text{ cm}^{-3}$ ) because of the short baseline (4 m). The FAST mission, however, was designed such that its spin plane is often nearly coplanar (within  $6^\circ$ ) of the plane formed by the magnetic field and the spacecraft velocity, so the instantaneous parallel electric field does not greatly depend on the short baseline measurement. At the time of the event, the magnetic field was  $2.7^\circ$  from the spin plane. We conclude that the  $200 \text{ mV/m}$  spike in  $E_{\parallel}$  was measured with an accuracy of  $\pm 20 \text{ mV/m}$ .

Figure 3 displays a cartoon interpreting the observations in Fig. 1. The thick, vertical solid line represents the magnetic field and the thin, “U”-shaped solid lines represent equipo-

tential contours of a two-dimensional structure that carries a parallel electric field. The dashed lines indicate the velocity of the spacecraft, the velocity of the structure, and the relative path of the spacecraft through the structure. The spacecraft entered the structure at  $\sim 15:35:03 \text{ UT}$  when the electric field was perpendicular to  $\mathbf{B}$  as attested by the negative spike in  $E_{\perp}$  and the appearance of anti-earthward electron fluxes. Although the actual spacecraft velocity was nearly perpendicular to  $\mathbf{B}$ , the apparent path of the spacecraft relative to the structure was at an oblique angle due to the motion of the structure as shown in Fig. 3. The spacecraft then exited the potential structure 2 s later at  $\sim 15:35:05 \text{ UT}$  through a region which had a parallel electric field. Thus, the spacecraft crossed the parallel electric field from a region of accelerated electrons (high-potential side) into the region of unaccelerated electrons (low-potential side).

## 2.2 Wave observations

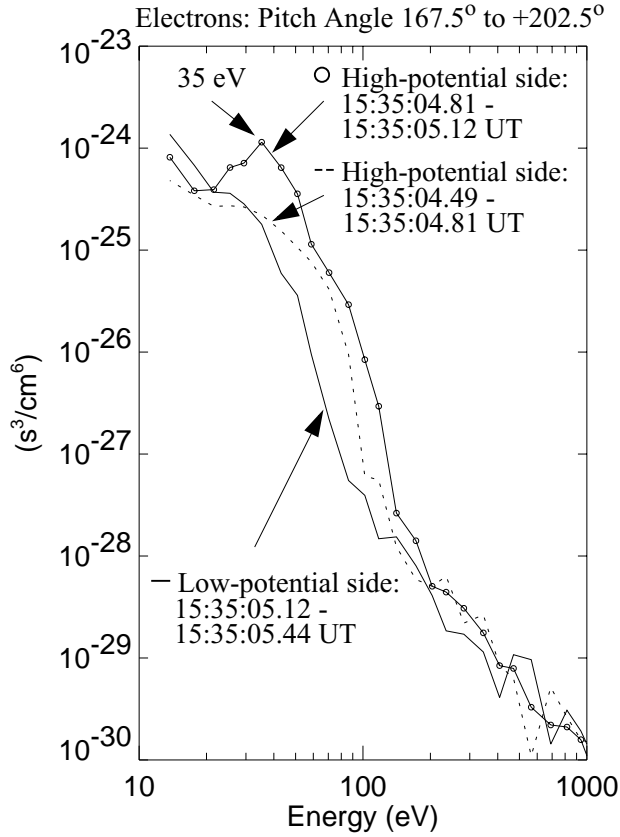
Figure 4 displays (a) an expanded view of  $E_{\parallel}$ , (b) the potential ( $\Phi_{\parallel}$ ) derived from integrating  $E_{\parallel}$ , and (c) the broad band electric field wave amplitude from 1 kHz to 16 kHz. The potential ramp was calculated with the assumption that the relative velocity between the spacecraft and the parallel electric field structure was 10 km/s. The net potential is  $\sim 27 \text{ V}$ .

A salient feature of the figure is the substantial increase in the wave power that was observed clearly before the parallel electric field. There was no corresponding increase magnetic field wave power, and hence we conclude the wave emissions were primarily electrostatic. The most intense wave power occurs over a  $\sim 50 \text{ ms}$  period. Under the interpretation diagrammed in Fig. 3, the intense wave emissions are on the high-potential side of the parallel electric field and associated with the accelerated electron beam. The energy density,  $W = \epsilon_0 E^2 / 2n_e T_e \cong 10^{-3}$  to  $10^{-2}$  (ignoring the magnetospheric electrons), indicates strong modification of the electron distribution and possibly nonlinear behavior. Far less wave energy is seen on the low potential side of the double layer.

The intense wave emissions were broad banded with the frequency extending up to and possibly over 16 kHz (Fig. 1e). The plasma frequency was  $\sim 40 \text{ kHz}$ . No waveform data are available for this event, but Ergun et al. (2001) found that the spectral signature of the wave emissions was that of electron phase-space holes. This interpretation has since been verified by numerical simulations (Newman et al., 2001) and subsequent observations (Andersson et al., 2002). The electron phase-space holes in the downward current region of the aurora were seen to travel with the electron beam and thus explain the wave turbulence beginning at  $\sim 15:35:02 \text{ UT}$ .

## 2.3 Electron observations

Figure 5 displays the measured electron distributions on the low-potential side (solid line) and high-potential side (circles and dashed lines) of the parallel electric field. In all of

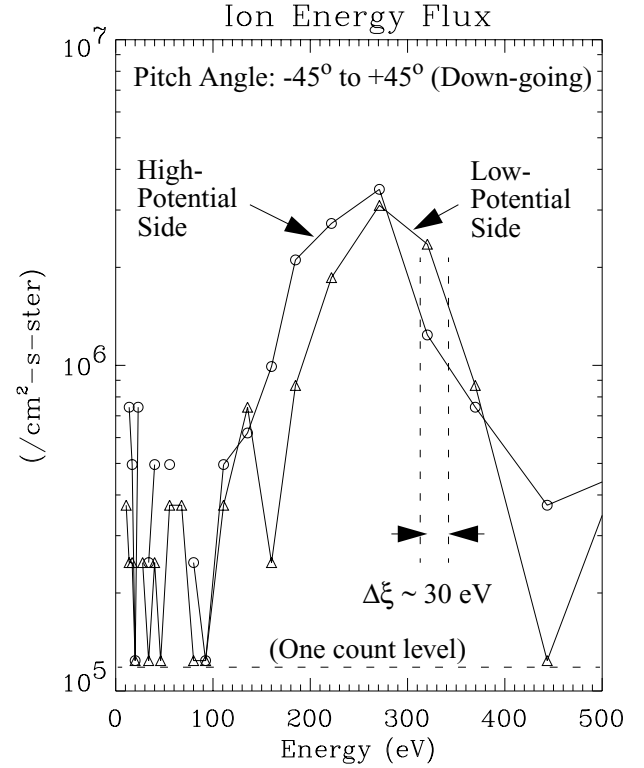


**Fig. 5.** Electron distributions surrounding the parallel electric field. The solid line is the distribution on the low-potential side. The circles indicate the distribution closest to the beam region. An unstable peak is visible. The dashed line is the distribution on the high potential side farther from the double layer.

the distributions, there is a tenuous,  $\sim 1 \text{ cm}^{-3}$ , hot,  $\sim 1.2 \text{ keV}$ , electron population indicating a magnetic connection to the central plasma sheet that is not apparent in Fig. 5, but can be seen as the blue background in Fig. 1. This population is not significantly perturbed by the parallel electric field.

The low-potential or source distribution (Fig. 5, solid line) displays a heated population with a density  $n_e \sim 5 \text{ cm}^{-3}$  and a parallel temperature  $T_e \sim 6 \text{ eV}$  (derived from the second moment of the electron distribution in the energy range from 5 eV to 100 eV) that is drifting anti-earthward at  $\sim 1 \text{ km/s}$ . Langmuir probe measurements (not displayed), however, suggest that the source electron distribution has a dense ( $n_e \sim 50 \text{ cm}^{-3} \pm$  a factor of two) core with a temperature that is below the energy range of the electron analyzer ( $< 5 \text{ eV}$ ). Thus, the electron distribution on the low potential side is comprised of three populations: a cold, dense core, a “heated” ( $T_e \sim 6 \text{ eV}$ ) population and a “hot” ( $\sim 1 \text{ keV}$ ), tenuous plasma-sheet population. The second moment of the three combined populations yields an electron temperature of  $\sim 25 \text{ eV}$  which indicates that the ion acoustic speed in an  $\text{O}^+$  plasma is roughly  $12 \text{ km/s}$  (within a factor of 2).

The high-potential or accelerated electron distribution (cir-



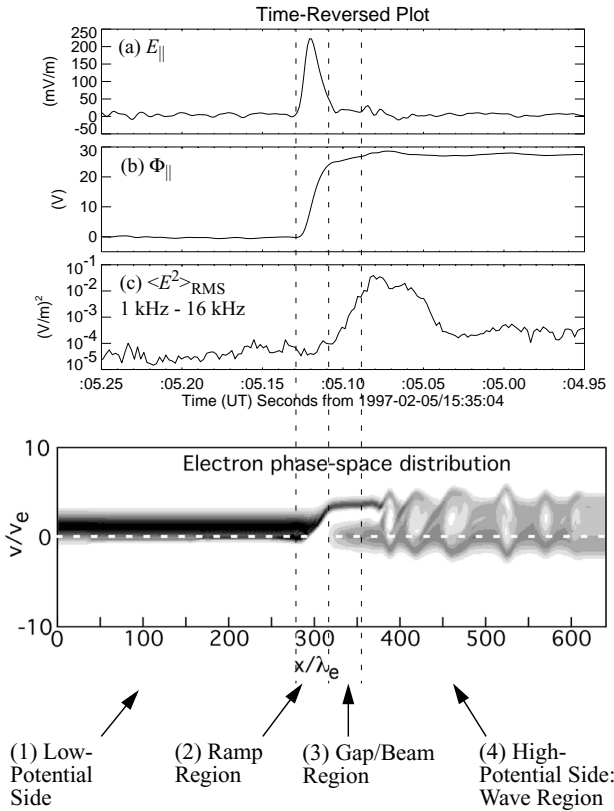
**Fig. 6.** Ion energy fluxes measured on both sides of the parallel electric field. The circles represent the fluxes on the high-potential side and the triangles represent the fluxes on the low potential side. The ion fluxes are consistent with a 30 V potential. The one count level is indicated on the plot. Measured flux levels under  $\sim 10^6 \text{ (cm}^2\text{-s-ster)}^{-1}$  have large statistical uncertainties.

cles) shows a dramatic increase in fluxes with a peak at  $\xi = 35 \pm 4 \text{ eV}$  in energy, indicating a net energy gain ( $\Delta\xi$ ) of roughly 30 eV. The moments of the distribution in the energy range from 5 eV to 100 eV yield a parallel temperature of  $\sim 12 \text{ eV}$  and a density of  $12 \text{ cm}^{-3}$  (accurate to a factor of two). The perpendicular temperature is less than 2 eV. Within error,  $\Delta\xi$  agrees with the net potential ( $\Delta\Phi = 21 \text{ eV}$  to  $30 \text{ eV}$ ) derived from the electric field observations. This agreement strongly supports the interpretation diagrammed in Fig. 3.

The electron distribution farther from the parallel electric field on the high-potential side (Fig. 5, dashed line) has no peak, rather, it appears as if heated and drifting. The accelerated electron distribution has a significant phase-space density with a positive parallel velocity (see Fig. 4c, Ergun et al., 2001) indicating that part of the population is travelling back toward the double layer. This electron population, created by the wave turbulence, may affect the structure and dynamics of the double layer.

## 2.4 Ion observations

The ion energy fluxes also indicate several different populations. Energy fluxes at pitch angles from  $-22.5^\circ$  to  $22.5^\circ$  (travelling earthward) are displayed in Fig. 1d. The most



**Fig. 7.** The structure of the parallel electric field region. The top panels are identical to those in Fig. 4 with the time axis reversed. The bottom panel is a snapshot of the electron phase-space from an numerical simulation (Newman et al., 2001). In this figure, the horizontal axis ( $x$ ) is the distance anti-parallel to the magnetic field in units of Debye lengths (derived from the low-potential side), increasing from low-altitude (low potential side) to high-altitude (high-potential side). The ramp region, the beam region and the turbulent region are seen in both observation and simulation.

energetic fluxes ( $>1$  keV) indicate a hot ( $\sim 5$  keV) plasma sheet population. Not visible in Fig. 1d are weak ion fluxes ( $\sim 3 \times 10^5 \text{ cm}^{-2} \text{ s}^{-1} \text{ sr}^{-1}$ ) at energies on the order of 10 eV with nearly perpendicular pitch angles. The perpendicular fluxes were detected after 15:35:05 UT indicating ion conics on the low-potential side of the double layer. The ion distributions associated with double layers are discussed by Andersson et al. (2002).

The most unusual feature in Fig. 1d is the enhancement at  $\sim 250$  eV that is visible throughout the plot. We examine these fluxes in more detail in Fig. 6 which plots the ion energy flux on the high- and low-potential sides of the double layer. One can see a clear peak in energy flux at  $\sim 250$  eV that was also visible in Fig. 1d. We interpret these fluxes as evidence of an approximately 250 eV parallel electric field that is anti-earthward of the spacecraft. Under this interpretation, the observed parallel electric field carries only part of the total potential on the flux tube. Figure 3 includes a U-shaped potential structure, anti-earthward of the satellite path, to ac-

count for the ion observations.

The horizontal axis in Fig. 6 is linear in energy which allows one easily to see a  $\sim 30$  eV increase in energy between the high-potential side and the low-potential side. The one-count level of the instrument, indicated in the figure by a dashed line, is such that values under  $10^6 (\text{cm-s-ster})^{-1}$  have large statistical uncertainties. Thus, the earthward-travelling ions see the same energy gain as the anti-earthward-travelling electrons, both of which are consistent with the observed parallel electric fields.

### 3 Discussion

The observed parallel electric field structure reaches a relatively high amplitude (200 mV/m) and endures for  $\sim 10$  to  $\sim 20$  ms (Fig. 4). Using 10 km/s as relative speed along  $\mathbf{B}$  between the electric field structure and the spacecraft, the size of the structure is between 100 and 200 m. The plasma parameters on the low-potential side of the parallel electric field are such that  $\lambda_D \sim 5$  m if one considers all electron populations (the magnetospheric electrons increase  $\lambda_D$ ). On the high-potential side,  $\lambda_D \sim 20$  m. Thus, the ramp region extends between 5 and  $40 \lambda_D$  along  $\mathbf{B}$ . The local ion-acoustic speed was estimated to be  $\sim 12$  km/s on the low-potential side and  $\sim 23$  km/s on the high-potential side (Ergun et al., 2001). The scale size and the speed of the observed electric field structure are consistent with the parameters of a strong double layer (Block, 1972).

A numerical simulation, designed to investigate the observed parallel electric fields illustrates the basic structure of the ramp, beam and turbulence regions. The details of the simulation are described elsewhere (Newman et al., 2001) so we provide only a very brief description. The simulation used a 1-D, open boundary Vlasov code with dynamic electrons. Figure 7 (bottom panel) displays a snapshot of an electron distribution near the end of a the simulation that started with a density depression in the presence of a current imposed through boundary injection. The simulation has ions drifting to the left at the ion thermal speed and the electrons are drifting to the right. Thus, Fig. 7 is in the rest frame of the double layer, that is, moving at the ion acoustic speed.

Figure 4 is replotted at the top of Fig. 7 with the horizontal axis reversed so electrons are accelerated from the left toward the right as in the numerical simulation. The evolved state of the numerical simulation is qualitatively similar to the observations. The regions of interest are marked on the plot with vertical dashed lines. On the low-potential (left) side, the electrons are undisturbed and drifting through the double layer. Ions are accelerated into the low-potential side through the double layer but cause only low-level turbulence. The ramp region is confined to roughly  $10 \lambda_D$ , carries a potential ( $\Delta\Phi$ )  $\sim 3$ – $5$  times the initial electron thermal energy, and appears nearly monotonic. Immediately to the right of the ramp (high-potential side) is an unstable electron beam. The convective instability can exist indefinitely but extends over a distance less than  $10 \lambda_D$ . To the right of the beam re-

gion, the beam is rapidly stabilized by intense turbulence and electron phase-space holes. The electron phase-space holes propagate toward the right, away from the double layer in the same direction as the accelerated electrons.

Although the double layer and intense electrostatic turbulence are spatially separated, they can interact by means of the stagnating electrons. Both the observations and the numerical simulations indicate that the intense wave turbulence creates an electron population that is moving toward the double layer. This population, in turn, is reflected by the double layer back into the wave turbulence region. The stagnating electrons appear to set up the two-stream instability which can lead to rapid growth of electron phase-space holes (Goldman et al., 1998). This population also must be considered in a self-consistent solution of the double layer.

The above observations suggest that multiple double layers may lie on a magnetic flux tube. The electrons on the low-potential side of the double layer appear to be a relatively undisturbed ionospheric population while the down-going ions indicate a parallel potential of  $\sim 250$  V on the same flux tube but at higher altitude. The numerical simulations and the observations indicate an electron acceleration of approximately 3 to 5 times the initial electron thermal velocity ( $v_e$ ), as measured on the low-potential side. Thus a second strong double layer may be responsible for the ion acceleration.

The concept of multiple double layers gets strong support from the observations reported by Andersson et al. (2002). In those double layer observations, the low-potential electron population was drifting into the double layer with  $\sim 50$  eV initial temperature before experiencing a 700 V acceleration. The source electrons appear to have been accelerated by a double layer earthward of the spacecraft.

#### 4 Conclusions

We presented direct observations of the parallel electric field in the downward current region of the auroral zone. The observations are consistent with a strong double layer moving along  $\mathbf{B}$  at the ion acoustic speed in the same direction of the accelerated electrons. The potential drop extends  $\sim 10 \lambda_D$  along  $\mathbf{B}$ . Intense electrostatic emissions are spatially separated from the structure on the high-potential side. Electron phase-space holes emerge from the wave turbulence associated with the double layer.

The potential structure accelerates electrons to several times their initial thermal velocity which results in a factor of 10 gain from the initial thermal energy. Intense quasi-electrostatic wave emissions and electron phase-space holes rapidly modify the accelerated electron distribution. Part of the electron distribution (stagnating electrons) is reflected back into the double layer through interaction with the intense wave turbulence. Thus, the intense wave turbulence may interact with the double layer through this stagnating electron population.

Ions are accelerated into a beam in travelling in the opposite direction of the electron beam. The energetic ion fluxes suggest that parallel electric fields on a magnetic flux tube are a localized to one or several strong double layers. There are, however, several open questions that need to be resolved including how the electrostatic turbulence and double layers interact through the stagnating electrons, how moving double layers interact with heated ions, the role of the accelerated ions and if they disrupt the double layer, and how several moving double layers combine to form the large-scale, quasi-static potential structure.

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